

Progress on radiative transfer modeling in edge plasmas

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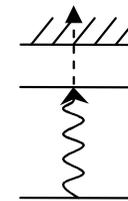
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Introduction

Tokamak edge and divertor plasmas can be optically thick to the neutral hydrogen radiation

The photon absorption provides a source of excited atoms
They can ionize more easily



“Photo-induced”
ionization

Consequences for ITER?

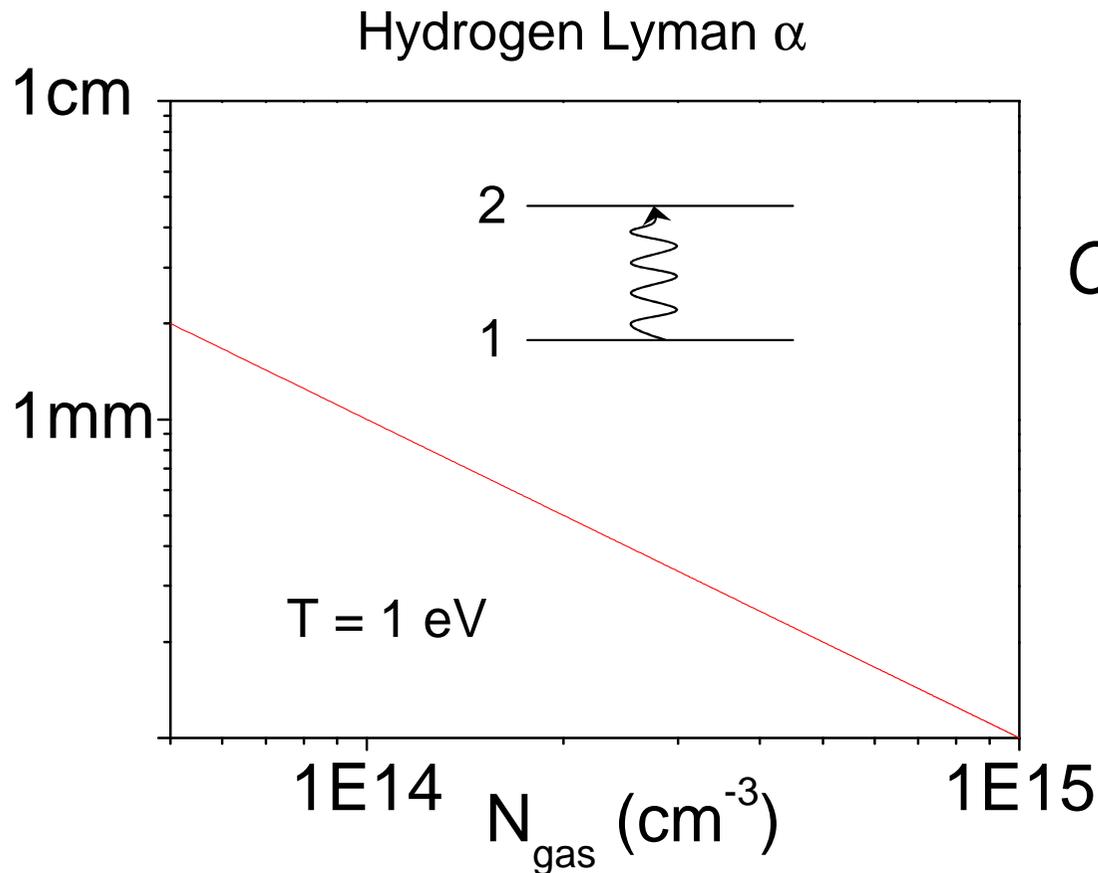
Efforts in theory and modeling are currently ongoing

- 1) Opacity effects in magnetic fusion
- 2) A transport model for radiation in fluctuating medium

Photon mean free path estimates

monochromatic mean free path $\propto \frac{1}{N_{gas}}$

Mihalas, Stellar Atmospheres



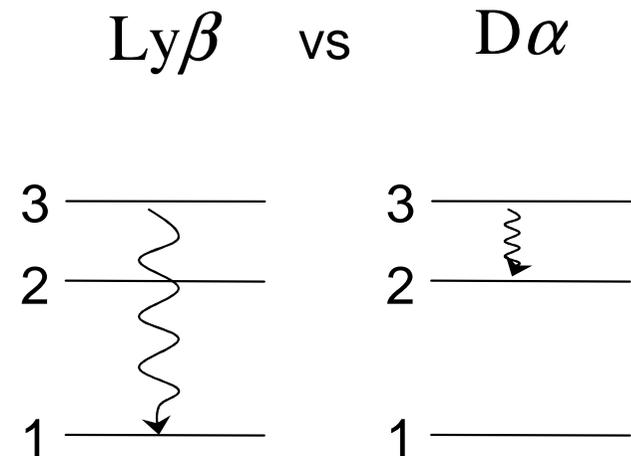
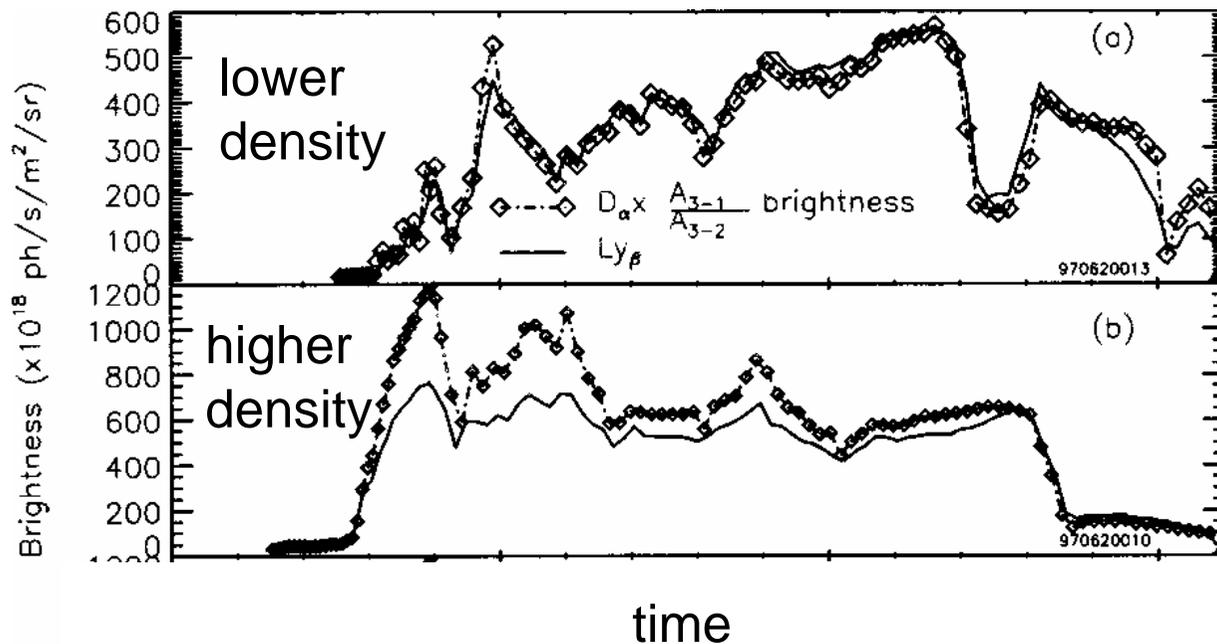
Opacity: large $N \times L$ machines

ITER & DEMO

Experimental observations: Ly β /D α

Ratio Ly β / D α in Alcator C-Mod: **proof of opacity**
in high-density divertors

J. L. Terry et al., PoP (1998)

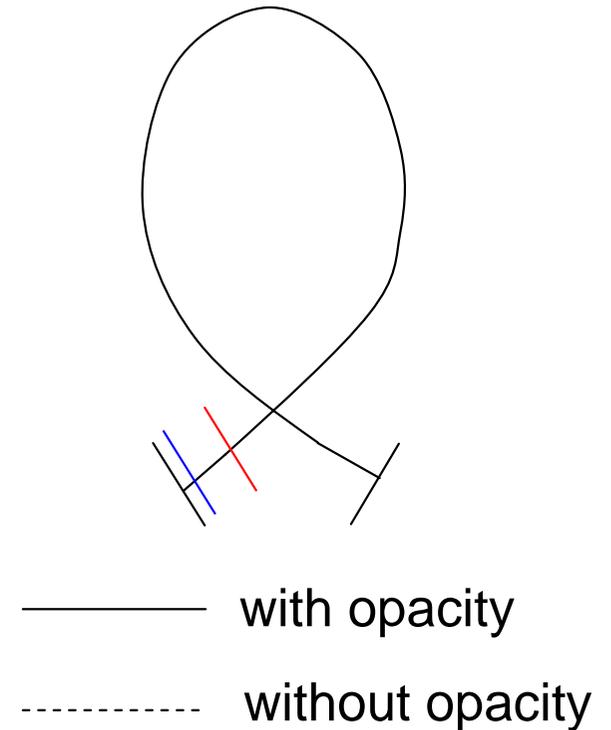
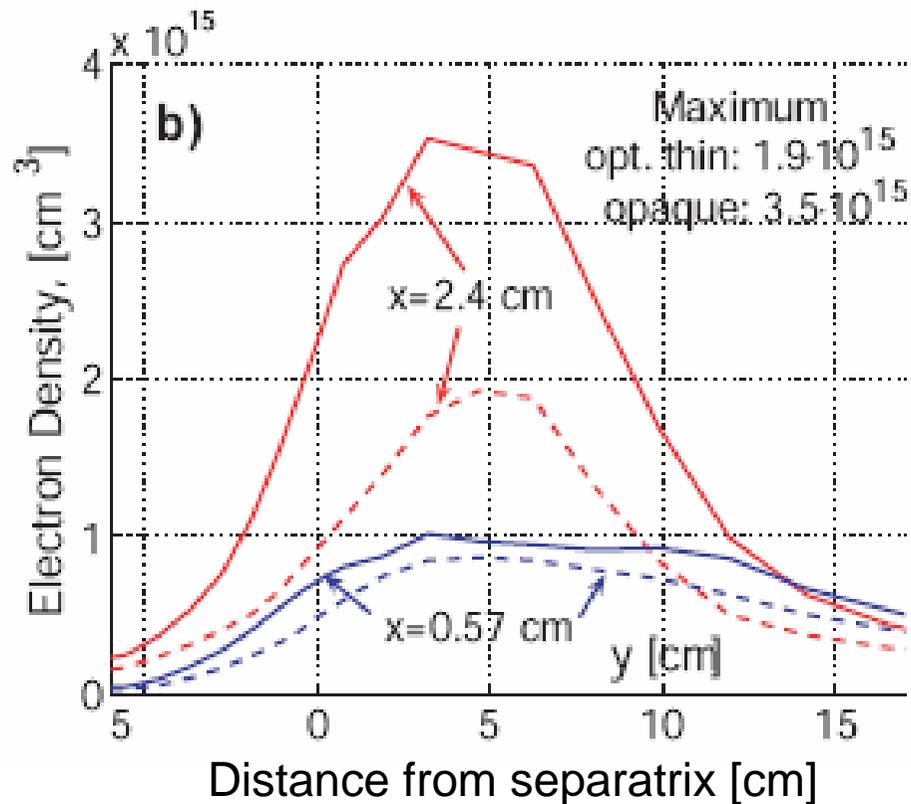


Similar observations at JET, but somewhat weaker

Numerical simulations: ITER

B2-EIRENE code (www.eirene.de)

Kotov, Reiter, Kukushkin et al., CPP (2006)



The electron density profile in the divertor is affected by line-radiation opacity

The transport model

- i) D-atoms in their fundamental state ($n = 1$):
Boltzmann equation by Monte-carlo method

- ii) Excited D-atoms ($n > 1$): stationary collisional-radiative model

- iii) The charged particles: fluid model (B2 code)

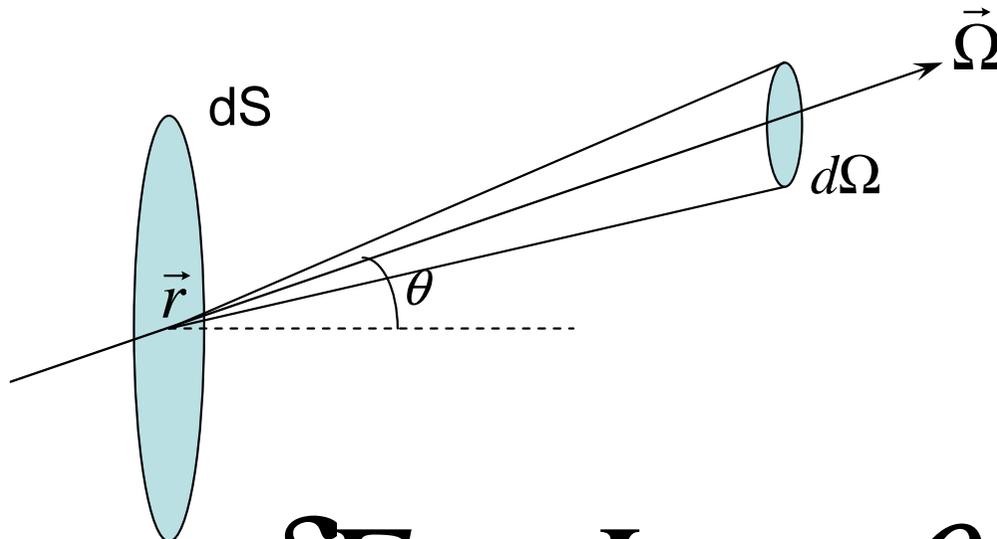
- iv) Photons: kinetic approach, Monte-Carlo method

Radiative transfer formalism

Topic associated with astrophysics, ICF

Fundamental quantity of interest = “radiation specific intensity”

$$I(\omega, \vec{\Omega}, \vec{r}, t)$$



$$\delta E = I \cos \theta d\omega d\Omega dS dt$$

The equation of transfer

A linear, Boltzmann-like transport equation

$$\frac{1}{c} \frac{\partial I}{\partial t} + \vec{\Omega} \cdot \frac{\partial I}{\partial \vec{r}} + \chi I = \eta$$

χ absorption

η spontaneous emission

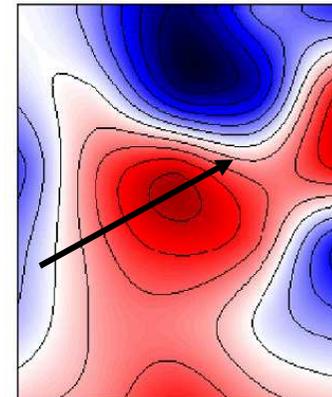
Not retained here:

- Stimulated emission (low radiation intensity, no pop. inversion)
- Scattering (complete redistribution)

Plasma fluctuations

All of the numerical investigations done so far were made assuming a plasma background whose typical variation scales are much larger than the neutrals & radiation transport scales

This is questionable
for tokamaks: $l_{\text{turb}} < 1 \text{ cm}$



Statistical approaches have been proposed recently for neutrals
Y. Marandet et al., PET 2009; PPCF (2011)
A. Mekkaoui et al., P1-6

Radiation transport?

Setting-up the formalism

i) Coarse-graining in time

$$I(\omega, \vec{\Omega}, \vec{r}, t) \rightarrow \langle I \rangle(\omega, \vec{\Omega}, \vec{r}, t) = \frac{1}{T} \int_t^{t+T} dt' I(\omega, \vec{\Omega}, \vec{r}, t')$$

ii) Decomposition

$$I = \langle I \rangle + \delta I \quad \chi = \langle \chi \rangle + \delta \chi \quad \eta = \langle \eta \rangle + \delta \eta$$

A quasilinear-type model

G. C. Pomraning,

Linear Kinetic Theory and Particle Transport in Stochastic Mixtures (1991)

The model is suitable for regimes where $l_{\text{turb}} \ll l_{\text{mfp}} \sim \langle \chi \rangle^{-1}$

$$\left(\vec{\Omega} \cdot \vec{\nabla} + \langle \chi \rangle \right) \langle I \rangle = \langle \eta \rangle - \langle \delta \chi \delta I \rangle$$

$$\left(\vec{\Omega} \cdot \vec{\nabla} + \langle \chi \rangle \right) \delta I = \delta \eta - \delta \chi \langle I \rangle + \langle \delta \chi \delta I \rangle - \delta \chi \delta I$$

$$\delta I = \left(\vec{\Omega} \cdot \vec{\nabla} + \langle \chi \rangle \right)^{-1} (\delta \eta - \delta \chi \langle I \rangle)$$

Coarse-grained transport equation

$$\left(\vec{\Omega} \cdot \vec{\nabla} + \chi_{eff}\right) \langle I \rangle = \eta_{eff}$$

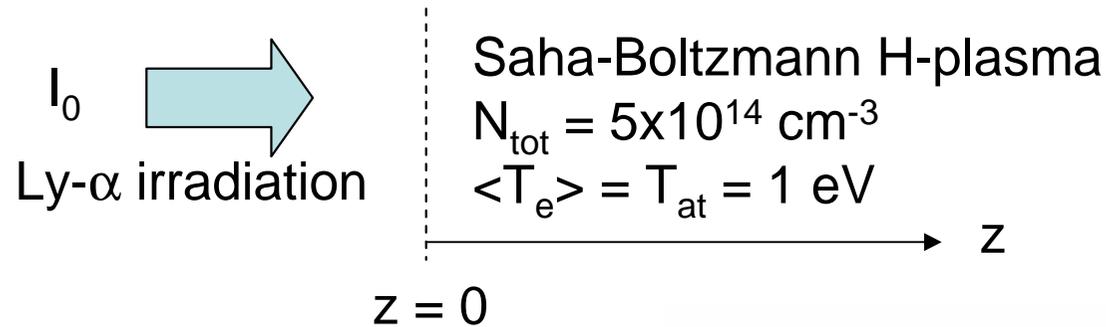
$$\chi_{eff} = \langle \chi \rangle - \int_0^\infty ds \langle \delta\chi(\vec{r}) \delta\chi(\vec{r} - \vec{\Omega}s) \rangle e^{-\langle \chi \rangle s}$$

$$\eta_{eff} = \langle \eta \rangle - \int_0^\infty ds \langle \delta\chi(\vec{r}) \delta\eta(\vec{r} - \vec{\Omega}s) \rangle e^{-\langle \chi \rangle s}$$

Homogeneous case: $\langle \chi \rangle$, $\langle \eta \rangle$ are space independent

In the absence of fluctuations,
the standard radiative transfer equation is recovered

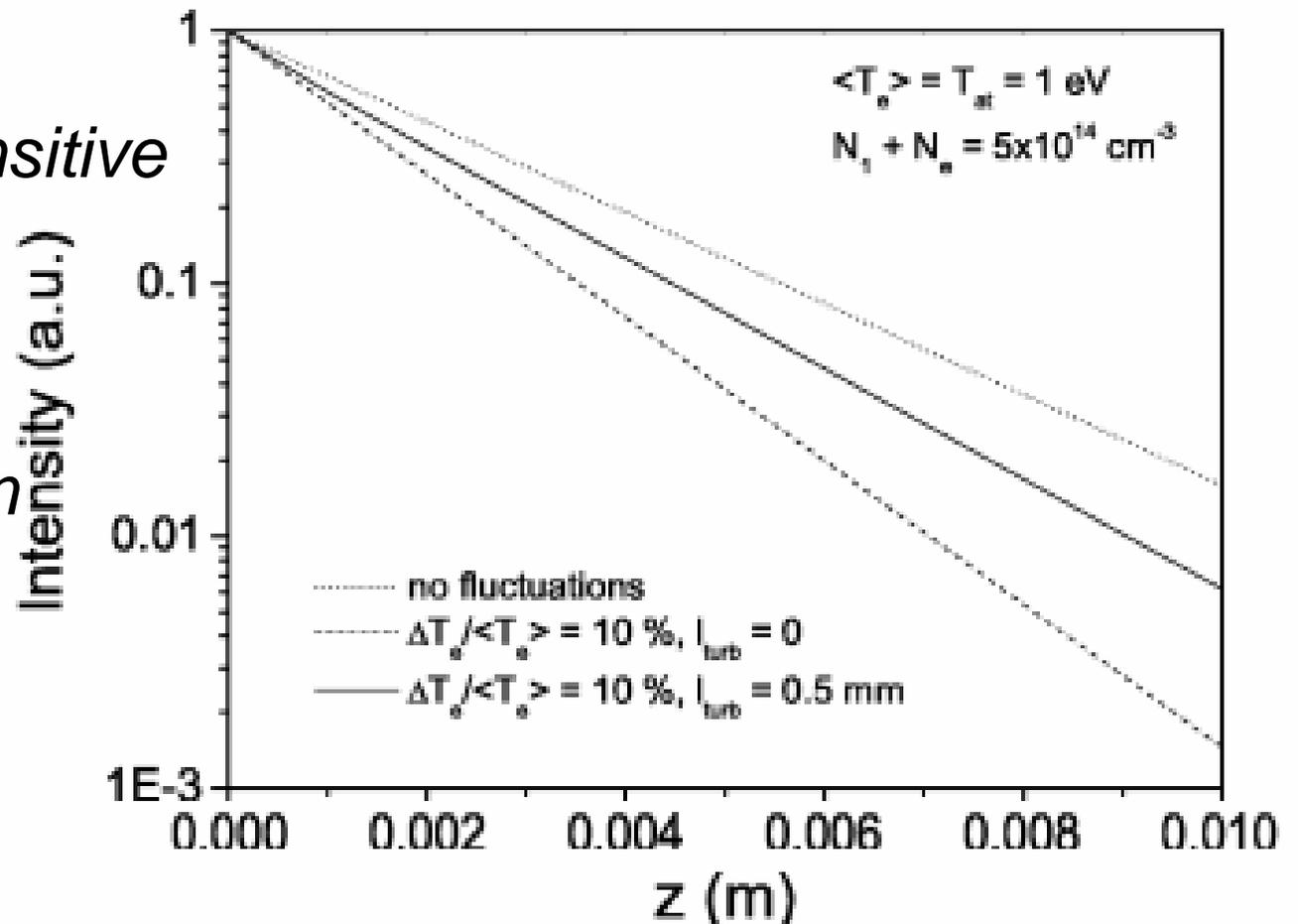
Attenuation of a radiation pencil



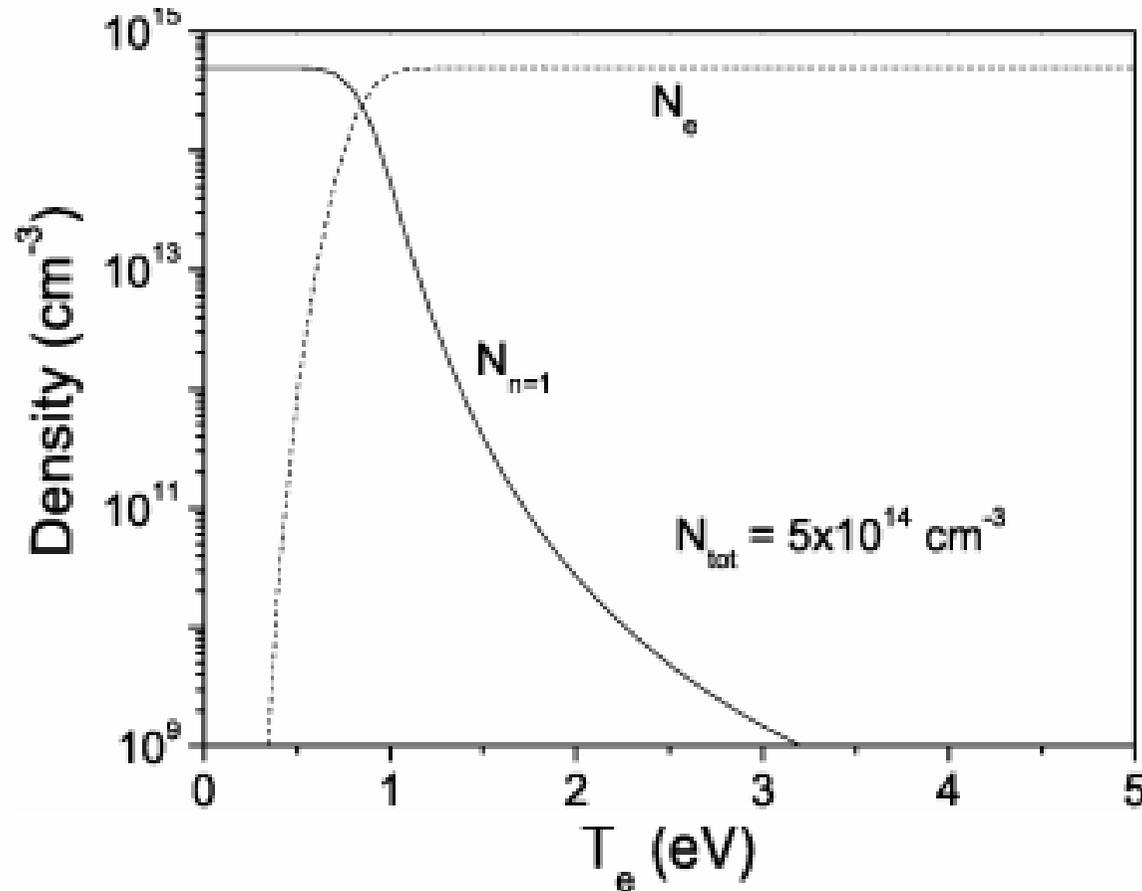
- Gamma PDF for T_e -fluctuations
- Exponential correlation
- Doppler line shape

i) The penetration is sensitive to the plasma fluctuations

ii) The spatial correlation reduces the global opacity



Why is the penetration sensitive to fluctuations?



$$\chi \propto N_{n=1}$$

The absorption coefficient χ is strongly dependent on T_e

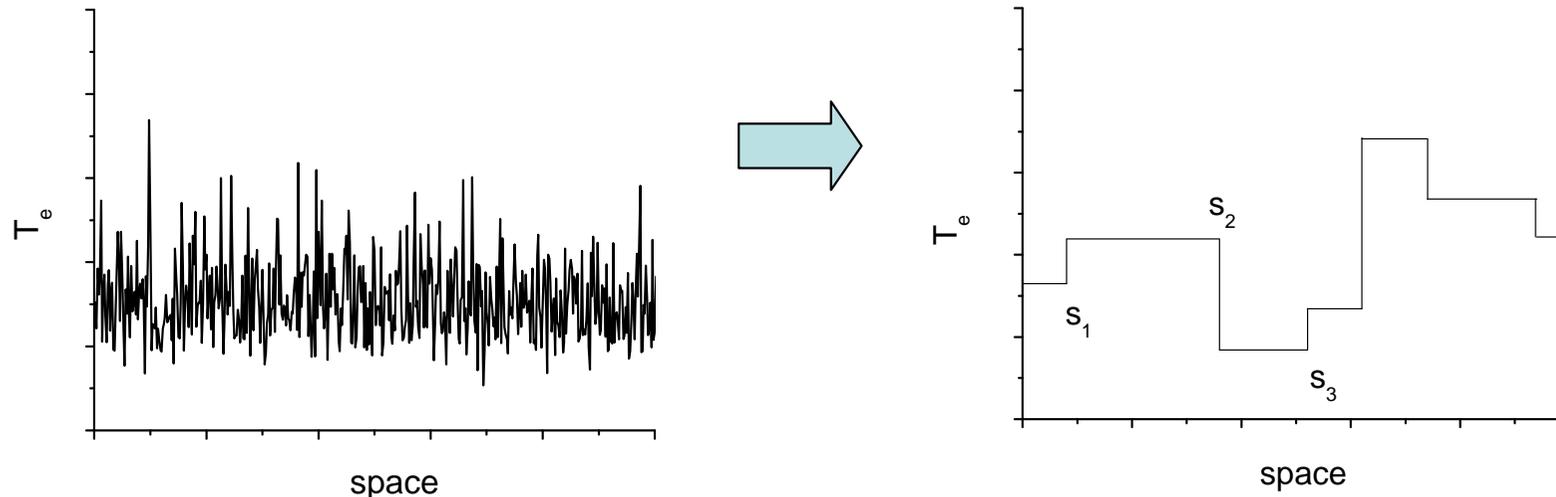
$$\langle \chi(T_e) \rangle \neq \chi(\langle T_e \rangle)$$

Similar result for the emission coefficient

Beyond the small $l_{\text{turb}}/l_{\text{mfp}}$ limit?

“The method of model coefficients”

A. Brissaud and U. Frisch,
J. Math. Phys. (1974)



- Same result as from the quasilinear theory (QLT) in the limit $l_{\text{turb}} \rightarrow 0$
- As in the QLT, use of the 1-point PDF and the spatial correlation function

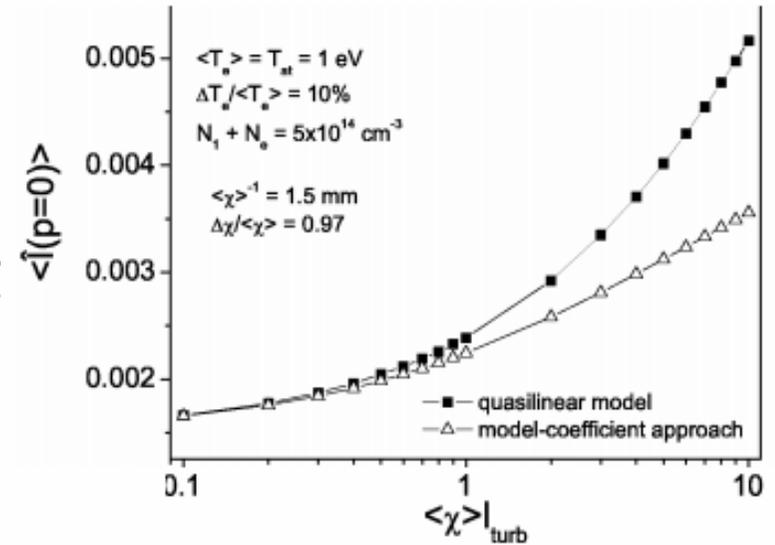
Laplace transform

$$\langle \tilde{I}(p) \rangle = \frac{\langle (p + 1/l_{\text{turb}} + \chi)^{-1} \rangle}{1 - \frac{1}{l_{\text{turb}}} \langle (p + 1/l_{\text{turb}} + \chi)^{-1} \rangle} I_0$$

Behavior at large z

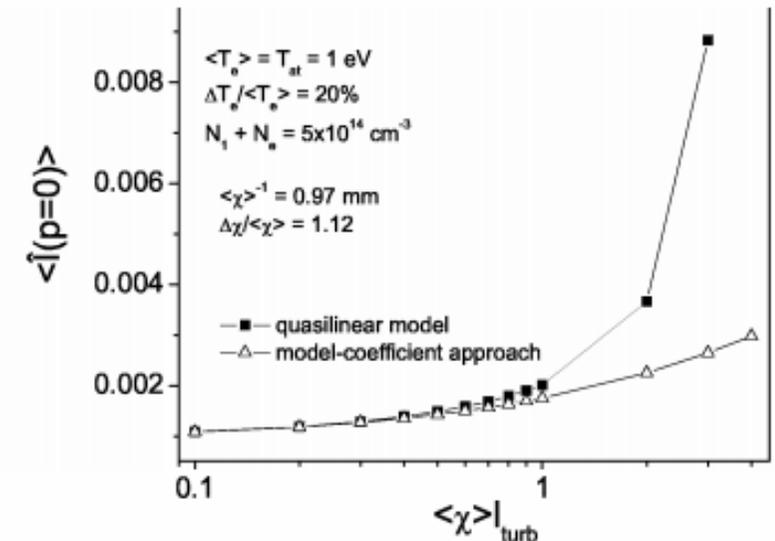
i) $\langle \tilde{I}(p=0) \rangle$ at $\Delta T_e / \langle T_e \rangle = 10\%$

The QLT is in a rather good agreement with the model-coefficient approach (max. ~40 % deviation)



ii) $\langle \tilde{I}(p=0) \rangle$ at $\Delta T_e / \langle T_e \rangle = 20\%$

Deviations are much stronger
 This is a feature of large fluctuations:
 here $\Delta \chi / \chi > 1$



Summary

Opacity in Lyman series is omnipresent in divertor plasmas with large $N_{\alpha L}$: Alcator C-Mod, ITER, commercial fusion reactor?

Line radiation trapping can be accounted for in edge simulations by a kinetic approach

The radiation field is affected by plasma fluctuations: here, illustration with a simple transport model

Self-consistent atom-photon simulations?